

#613

VOYAGER 2

SATURN TITAN DELTA SCORPII ENCOUNTER DATA

77-076A-11B

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SATURN, TITAN, DELTA SCORPII ENCOUNTER

77-076A-11B

THIS DATA SET HAS BEEN RESTORED. ORIGINALLY IT CONTAINED ONE 9-TRACK, 1600 BPI TAPE WRITTEN IN ASCII. THERE IS ONE RESTORED TAPE. THE DR TAPE IS A 3480 CARTRIDGE AND THE DS TAPE IS 9-TRACK, 6250 BPI. THE ORIGINAL TAPE WAS CREATED ON A VAX 11/780 COMPUTER AND WAS RESTORED ON THE MRS SYSTEM. THE DR AND DS NUMBER ALONG WITH THE CORRESPONDING D NUMBER AND TIME SPAN IS AS FOLLOWS:

DR#	DS#	D#	FILES	TIME SPAN
DR005286	DS005286	D060131	9	08/12/81 - 08/25/81

REQ. AGENT
DAD

RAND NO.
V0227

ACQ. AGENT
WSC

VOYAGER 2

SATURN, TITAN, DELTA SCORPII ENCOUNTER DATA

77-076A-11B

This data set consists of 1 tape. The tape contains 9 files.
The tape is 9-track, 1600 BPI, written in ASCII and was created on an
VAX 11/780 computer. The D and C numbers along with the time span follow:

<u>D#</u>	<u>C#</u>	<u>TIME SPAN</u>
D-60131	C-23639	08/12/81 - 08/25/81

LABORATORY FOR ATMOSPHERIC AND SPACE PHYSICS

VOYAGER 2 PHOTOPOLARIMETER

PROCESSED DATA FOR THE

NATIONAL SPACE SCIENCE DATA CENTER

The Voyager 2 Photopolarimeter (PPS) Saturn encounter data consists of atmospheric measurements for Saturn and Titan and a stellar occultation of Saturn's ring system. A description of the PPS instrumentation is given by Lillie et al. (1977) and the Voyager 2 preliminary Saturn Encounter results are reported in Lane et al. (1982).

The intensity and polarization measurements for Saturn and Titan were obtained using the 2640Å (300Å bandpass) and 7500Å (250Å bandpass) filters, the 135° and 45° analyzers, the 0.11 degree circular aperture (full width at half maximum), a 0.4 second integration period, and low-gain voltage. The PPS data points that correspond to one Voyager Television frame (80 integrations, or 48 seconds) are accumulated into one set of Intensity and Percent Polarization values for each of the two filters; the time resolution of each Intensity and Polarization pair is 11 seconds. Observations made during scan platform slewing are not included.

The Saturn data sets are discussed by West et al. (1983a). This paper presents the Saturn tables provided herein. The Saturn data are divided into three groups: 1) the "East-West Map" , 2) the "North-South" scans, and 3) the "Belt" and "Zone" scans. The "East-West" Map was a raster scan of the 0.11°

PPS field of view across Saturn's northern hemisphere at a phase angle of 10° . Several additional scans are included in this group. These "latitude scans" were limb-to-terminator scans at five latitude bands. Group two, the three "North-South" scans, are along the central meridian. The high resolution limb-to-terminator scans in the "Equitorial Zone" and "North Equitorial Belt" at phase angles up to 68° are in the third group.

The whole disk intensity and polarization observations of Titan cover phase angles from 3° to 154° . West et al. (1982b) discusses the analysis of the Titan data in terms of model scattering atmospheres and presents the tabular Titan data provided here.

The stellar occultation data were obtained using the 2640\AA filter, the 45° analyzer, the 1.0° circular aperture, a 0.0075 second integration period, and high-gain voltage. The occultation data presented here have been averaged into bins of 600 integrations. Successive integrations are separated by a radial distance of approximately 100m; 600 point bins are therefore approximately 60 Km wide. The star Delta Scorpii was observed continuously from its emersion from behind Saturn's disk through the shadowed D, C, B, A, and F rings. The discussion of the analysis of this 60 Km ring data is presented by Esposito et al. (1983).

Dr. Arthur L. Lane is the principal investigator for the Photopolarimeter experiment. He is resident at the Jet Propulsion Laboratory, Pasadena, California, 91109. Drs. Charles W. Hord, Larry W. Esposito, and Robert A. West are Co-Investigators and are resident at the University of Colorado, Boulder, Colorado, 80309. Drs. David L. Coffeen and Makiko Sato are also Co-Investigators and are resident at the Goddard Institute for Space Studies, 2880 Broadway, New York, New York, 10025.

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APPENDIX

The enclosed tape has the following tape characteristics.

- 1) 9-track, 1600 bpi. Written on a Digital TU77 drive.
- 2) ANSI STANDARD tape headers and End-of-File (EOF) structure (7-bit ASCII characters) as per VAX 11/780 system software.
There is one Volume Header, one File Header, two Record Headers, one EOF, the data record, and two EOFs for each of the three data files on this tape.
- 3) Physical data blocks are 2048 bytes long; data files are 7-bit ASCII records containing a four byte 'control' word followed by the ASCII bytes. The control word consists of two bytes of End-of-Text characters (hexadecimal=0) followed by a two byte logical record length value. The length refers to the

total number of bytes and includes the 4-byte control word.

Logical records are blocked into physical records and a hexadecimal value of 5E is used as fill from the end of the last logical record to the end of the physical record.

- 4) The Vax writes bytes onto a 9-track tape in the following order:

Vax word	0 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15	
Vax tape	<-----> <----->	
word written	2nd	1st

The resulting tape (up is the tape beginning direction)

8 9 10 11 12 13 14 15	Word part 2
0 1 2 3 4 5 6 7	Word part 1

- 4) An annotated dump is attached which shows the contents of the first five physical blocks. The left side of the dump shows the hexadecimal word (read from right to left) and the right side of the dump shows the ASCII equivalent word contents (read from left to right). The last column on the right is the hexadecimal 4-byte word number for the rightmost four bytes in the hexidecimal dump section.

\$ JOB 9:32:39
\$ SASS IN TWO
\$ NOP ** LIST OF EXECUTABLE FILES 1 2 3 4 5 6 ***

INPUT PARAMETERS ARE: AS AL

***** JCB DCNE •
\$EXE TPLIST ES

OBSERVATIONS OF SATURN C005 0103									
Voyager 2 Photopolarimeter 0005 0065									
FILE NO.	1	TAPE NO.	1	RECORD	1	LENGTH	2048	FILE NO.	1
2640A	750CA0101	Lat.	Lon.	a	CHI	MU			
1/F	P(%)	1/F	P(%)	1/F	East-West	North-South			
Equatorial Zone (0 to 15°)	Lat.	010101	3.3	76.2	10.8	17.7	0.412	0.566	175.8
0.312 -5.8	0.467	1.70101	6.1	73.4	10.7	17.7	0.470	0.616	175.3
1.1 0.300 -5.0	0.483	1.40101	5.1	78.0	10.7	17.7	0.403	0.556	175.6
2.18 0.348 -8.1	0.490	-0.10101	13.1	70.7	10.7	17.7	0.537	0.667	173.5
1.030 0.337 -3.6	0.484	1.50101	2.8	71.5	10.7	17.7	0.599	0.732	175.9
1.010 0.250 -4.0	0.581	1.80101	5.8	70.8	10.6	17.7	0.623	0.750	175.2
1.008 0.259 -3.6	0.591	1.40101	6.7	70.6	10.7	17.7	0.634	0.759	174.9
1.009 0.304 -3.6	0.596	1.10101	13.8	65.5	10.6	17.7	0.715	0.816	171.0
4.0 0.984 0.336 -2.3	0.545	0.20101	10.5	64.2	10.6	17.7	0.801	0.890	17.0
1.6 0.996 0.301 -0.1	0.629	0.10101	10.1	62.7	10.5	17.7	0.821	0.905	
171.4 0.996 0.293 -4.8	0.625	-0.10101	8.7	64.2	10.5	17.7	0.808	0.897	
172.8 1.002 0.296 -2.6	0.658	-0.10101	13.5	60.5	10.5	17.7	0.853	0.9	
23.167.3 0.971 0.335 2.6	0.578	0.30101	13.5	58.4	10.4	17.7	0.923	0	
0.969 159.9 0.966 0.295 0.1	0.631	0.50101	8.5	55.6	10.5	17.8	0.974		
0.995 180.0 1.000 0.301 1.4	0.665	0.50101	13.0	54.0	10.3	17.8	0.958		
5.0 0.996 101.0 0.961 0.314 1.2	0.609	0.00000							

0100 The first four columns give the planetocentric latitude and the System III longitude by Desch and Ong Kaiser (1981), the phase angle α , and the rotation angle CHI between the normal to the scattering plane, and the polarization axis of the analyzer. In the next three columns are Mu, the cosine of the emission angle, Mu0 the cosine of the incidence angle, and dPhi, the azimuthal angle in the scatterer's plane. The remaining columns give the blur correction factor for 2640 Å C₂T₂, and I/F and P for the two 0013 filters.

F 5 *** JOB DONE.

\$AVF IN 4
\$EXE TPLIST BS

INPUT PARAMETERS ARE: FILE 5 AS FL=1=1

TAPE NO. 1 FILE NO. 1
RECORD 0074 LENGTH 2048
WHOLE DISK OBSERVATIONS OF TITAN0124

a	b	c	d	e	f	g	h	i	j	k	l	m	n	P	O	S	CET	DT	NOTES0105 (D)	(%)	+01
0.6	24:1.46	18 (DEGREES)	<DN6>	*	<DN7>	*	B (10^6 Km)	CHI	A												-1.5
1.00106	226:10:41	3.8	6.7	0.1	6.7	0.1	10.9	12.90	5.1	0.045											
1.5	1.20106	227:12:41	2.0	2.8	6.8	0.1	6.9	0.1	0.4	12.00	-10.88	0.043									
0.48	-0.9	0.80115	228:12:08	10	2.7	8.7	0.1	8.8	0.1	0.8	11.40	-1.0	0.								
0.2	0.056	2.6	9.0	0.2	J-0FF0114	228:12:00	8	3.5	529.9	1.0	559.2	1.2	120.0	10.80	-1						
)	2.4	10.80	-10.2	0.046	3.0	1.7	J-0N0106	8	10.7	0.1	11.3	0.2									
)	2.5	0.5	12.1	10.00	-3.6	0.038	3.3	40106	20	5.4	20.0	0.1									
)	0.5	4.3	8.0	0.17	30.7	10.5	0.046	16.1	3.10125	13.52	8	12.0									
)	17.1	0.2	17.3	0.2	3.6	7.93	0.040	-0.8	1.6	BEFORE FILTER	00124	23									
)	1:14:10	20	12.0	21.3	0.3	20.4	0.3	3.6	7.93	23.0	0.050	-5.9	1.8								
)	AFTER FILTER	00106	232:13:30	18	16.2	21.4	0.1	22.2	0.2	0.3	6.73	29.2									
)	0.045	0.9	0.90106	233:10:17	20	3	31.8	0.1	33.1	0.2	33.0	0.2	0.0	5.5							
8	32.1	0.047	1.2	0.90106	233:12:49	20	3	31.3	0.1	31.3	0.2	33.0	0.2	0							
0	5.43	33.1	0.044	3.1	1.1	0.75	35.8	0.082	2.4	0.10106	237:19:55	10	154	16793.0							

TAPE NO.	FILE NO.	RECORD NO.	4 hour minute	LENGTH	2048	1																	
0106	23Z:01:10	9	7.8	23211.6	16.7	26025.1	6.5	40.0	0.82	38.8	0.120	45.6											
0	20120	237:01:20	10	78	35.9	0.0	41.0	0.0	0.0	0.82	38.8	0.000											
51.1	0.0	BACKGROUND0106	237:04:14	20	87	23541.0	12.5	27166.0	4.8														
0.74	39.0	0.102	50.8	0.20106	237:06:35	20	99	26436.0	12.9	29754.0	5.2												
0.0	0.69	39.0	0.096	48.2	0.20106	237:10:05	20	117	25243.2	10.4	28447.6												
)	4.2	0.0	0.67	38.4	0.086	44.3	0.10106	237:15:28	20	140	19651.0	9.1											
)	2114.0	6	4.4	0.6	0.75	35.8	0.082	2.4	0.10106	237:19:55	10	154	16793.0										
)	11.9	17475.0	4.8	0.0	0.91	31.4	0.102	10.7	0.10125														

The spacecraft Event Time (SCET) of the beginning of the observation is given 0085 as Day-0f-Year (D0Y), hour and minutes. The duration of the observation is DT. 0080 The parameter e is the phase angle. The average count rates per 0.4 second 0077 integration in analyzers 6 and 7 are, respectively $\langle DN6 \rangle$ and $\langle DN7 \rangle$. The 0052 uncertainty in these quantities is computed by 0069. S1 G=(1/N)(S0T(SUM cf i=1 tc (DNi- $\langle DN \rangle$)*2))•0079 The background count is given in the column under B. The Titan-spacecraft 0083 distance D is expressed in 10 to the 6 Km. The angle CHI is described in the 0084 referred text. The parameter A is the derived geometric albedo times the phase 0044 law, and P is the derived polarization.

\$JOB 9:35:00
 \$ASS IN TWO
 \$REW IN
 \$NOLIST OF FILES 3 AND 9 ***
)
 \$AVE IN 7
 \$EXE IPLIST ES

INPUT PARAMETERS ARE: AS FL=1=1

TAPE NO. 1		FILE NO. 1		LENGTH 2048		DISTANCE COUNT		DEPTH 0076		COUNT		OPTICAL 0075 KM	
RECORD	DEPTH	DISTANCE	COUNT	OPTICAL (KM)	RATE	DEPTHD	COUNT	DEPTH	0076	DEPTHD	COUNT	OPTICAL 0075	
37.8	0.010074	74742.	38.3	0.01	74611.	40.0	0.00	74676.			0.010074	0.010074	
4873.	37.0	75070.	38.5	0.010074	75136.	37.8	0.01	75005.			36.3	0.04	
38.2	0.010074	75267.	38.1	0.01	75464.	38.1	0.010074	75333.	38.3	0.010074	75		
399.	38.1	0.01	75596.	38.3	0.010074	75611.	38.4	0.01	75530.	38.0	0.01		
8.9	0.000074	75793.	38.7	0.00	75920.	34.2	0.060074	75858.	36.6	0.030074	759		
24.	34.6	0.06	76121.	32.4	0.090074	76187.	30.8	0.11	76055.	34.0	0.07		
0.	0.110074	76318.	35.3	0.05	76515.	37.4	0.020074	76384.	34.9	0.050074	7645		
0.	35.8	0.04	76647.	37.8	0.020074	76712.	37.4	0.02	76581.	37.8	0.02		
8	0.030074	76844.	36.5	0.03	76905.	36.1	0.040074	76778.	36.				
36.1	0.04	77172.	31.4	0.100074	77238.	30.1	0.120074	77106.	29.8	0.13	76975		
0.	0.060074	77369.	30.8	0.11	77435.	35.2	0.05	77303.	34.7				
34.8	0.05	77566.	35.2	0.050074	77632.	35.5	0.040074	77500.	37.5	0.01			
77657.	37.7	0.02											

TAPE NO. 2		FILE NO. 1		LENGTH 2048		DISTANCE COUNT		DEPTH 0076		COUNT		OPTICAL 0075 KM	
RECORD	DEPTH	DISTANCE	COUNT	OPTICAL (KM)	RATE	DEPTHD	COUNT	DEPTH	0076	DEPTHD	COUNT	OPTICAL 0075	
0074	140622.	39.2	0.00	140781.	39.0	0.000074	140835.	39.0	0.0000074	140728.	38.6		
40888.	38.8	0.000074	140941.	39.0	0.00	39.0	0.00	140994.	38.9	0.0000	1		
074	141047.	39.0	0.00	141206.	35.1	0.000074	141259.	38.6	0.0000074	141153.	38.5		
0.	0.00	0.00	141365.	38.4	0.01	141416.	38.7	0.00			14		
1312.	38.8	0.000074	141523.	38.6	0.000074	141576.	39.0	0.0000074	141416.	39.2	0.0000		
74.	141470.	38.7	0.00	141629.	33.0	0.000074	141682.	38.6	0.0000074	141576.	38.9		
0.	0.00	0.00	141787.	39.0	0.00	141840.	39.0	0.0000074	1414997.	38.5	0.01007	141	
734.	38.2	0.010074	142469.	38.7	0.00	141945.	39.2	0.0000074	141840.	38.5	0.01007	142	
4.	141892.	38.7	0.00	142050.	39.0	0.000074	142102.	38.9	0.0000074	141997.	38.8		
55.	39.0	0.000074	142207.	39.0	0.00	142260.	39.0	0.0000074	142260.	39.0	0.000074	1421	
0.	142312.	38.7	0.00	142365.	38.4	0.01	142417.	38.9	0.0000074	142417.	39.2		

**** JOB DONE.
 \$AVF IN 8
 \$SEX TPLIST BS

INPUT PARAMETERS ARE: AS AL
 TAPE NO. 1 FILE NO. 1

Dump of DEV253_NFT1: 8C 12-3110-1917 16-76-14-00

block number (0,0), 1